Observing Filaments Between Galaxy Clusters with NenuFAR How to Probe Cosmic Magnetic Fields

NenuFAR workshop presentation Etienne Bonnassieux, DRANOEL postdoctoral research fellow at UniBo

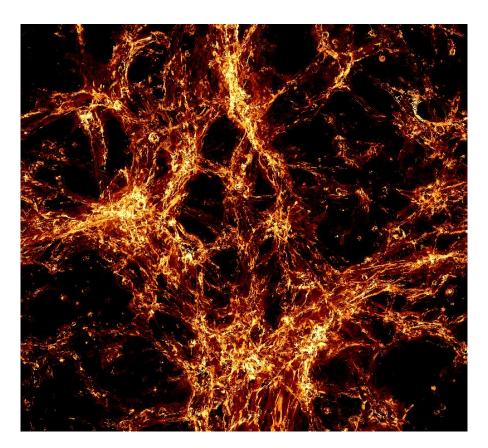
Collaboration with A. Bonafede, F. Vazza, C. Ferrari, F. de Gasperin, C. Tasse, ...

Outline

- Science Goals

- Why NenuFAR?

- Why us?



To study cosmic magnetism:

- Magnetic field in clusters dominated by ICM evolution
- To understand primordial magnetic field, need to go outside clusters

BUT

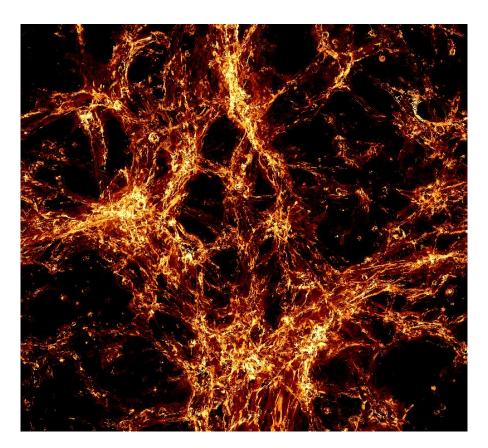
 If there are no relativistic electrons to decay in a forest, what can astronomers know about its magnetic field?...

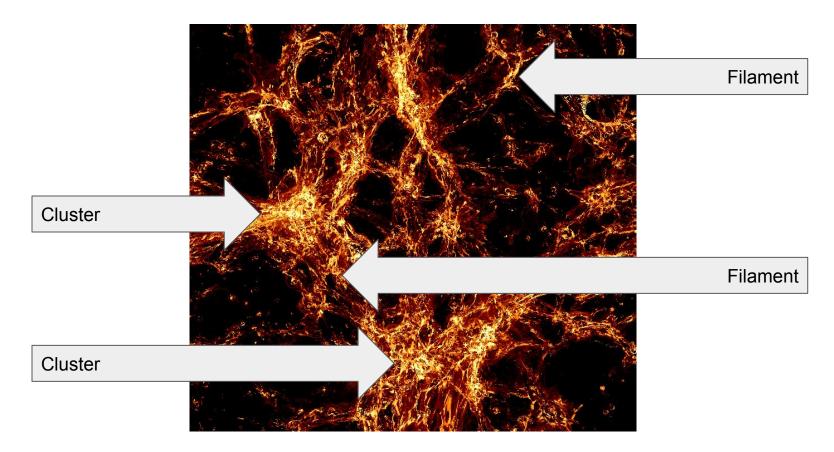
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Filaments: best environment for both presence of emitting particles and

unmarred cosmic magnetic field







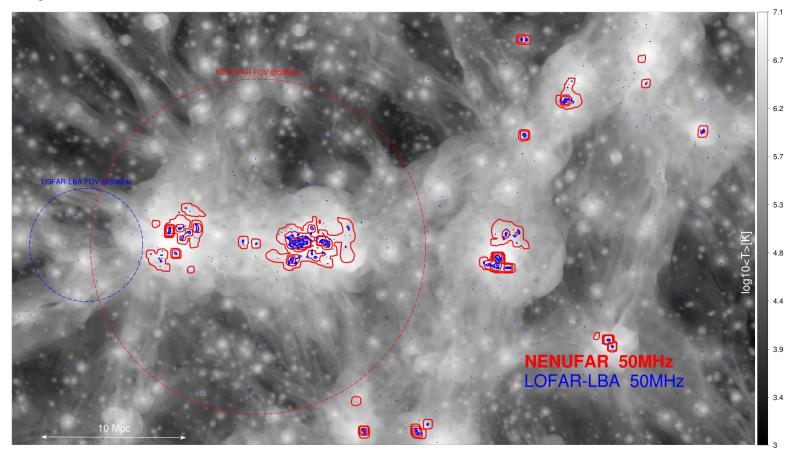
Observational constraints for filament observations:

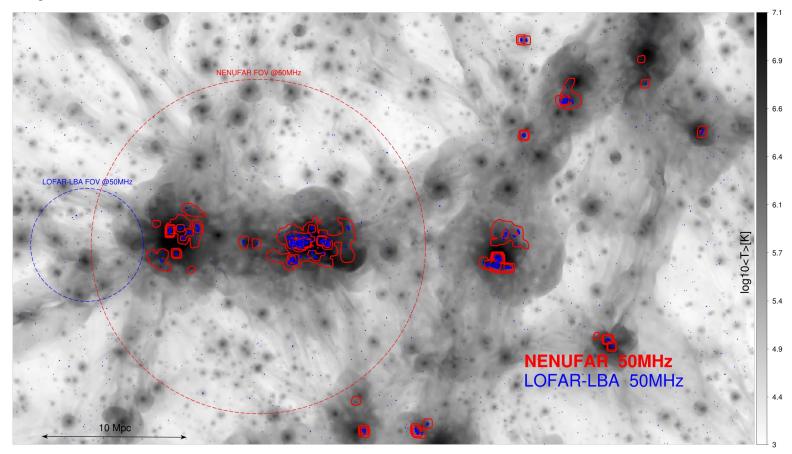
- Physics of cosmic magnetism traced by large-scale, diffuse structures
- Emission mechanism of said structures more effective at lower frequencies
- Presence of foreground/embedded point sources imposes angular resolution constraints to improve confusion limits

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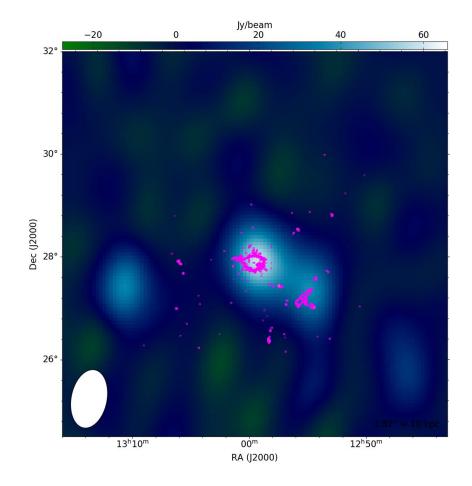
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NenuFAR standalone in imager mode + supersynthesis





Current Results





Article

First Lights on the Cosmic Magnetism Project with NenuFAR: the Coma Cluster field

Etienne Bonnassieux^{1,2}, Evangelia Tremou³, Julien Girard³, Alan Loh³, Valentina Vacca⁴ Philippe Zarka³, Jean-Mathias Griessmeier5,6, L. ("Leon") V.E. Koopmans7, Michel Tagger5, Stephane Corbel6,8, Gilles Theureau5,6

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- Universita di Bologna, Via Zamboni, 33, 40126 Bologna BO, Italy
- INAF. Instituto di Radioastronomia, Via Piero Gobetti
- 3 LESIA & USN, Observatoire de Paris, CNRS, PSL, SU/UP/UO, 92195 Meudon, France
- 4 INAF Osservatorio Astronomico di Cagliari, Via della Scienza 5, I-09047 Selargius (CA), Italy 5 LPC2E - Universite d'Orleans / CNRS, 45071 Orleans cedex 2, France
- Station de Radioastronomie de Nancay, Observatoire de Paris, PSL Research University, CNRS, Univ. Orleans, 18330 Nancay, France
- Kapteen Astronomical Institute, University of Groningen, P.O.Box 800, 9700AV Groningen, the Netherlands
- ⁸ AIM, CEA, CNRS, Université de Paris, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France

NO STREET NOT AND THE PROPERTY OF THE PROPERTY the Cosmic Filaments and Magnetism Pilot Survey is observing sources with the array as it is still under construction - with 57 (56 a core, 1 distant) out of a total planned 102 (96 core, 6 distant) mini-arrays online at the time of observation - to get a first look at the 4 low-frequency sky with NenuFAR. One of its targets is the Coma galaxy cluster: a well-known object, host of the prototype radio halo. . It also hosts other features of scientific import, including a radio relic, along with a bridge of emission connecting it with the halo. It is thus a well-studied object down to 144 MHz. In this paper, we show the first confirmed NenuFAR detection of the radio halo and radio 7 relic of the Coma cluster at 30 MHz, with associated intrinsic flux density estimates: we find an integrated flux value of 81 ± 15 Jy for the radio halo, and 38 ± 8 Jy for the radio relic. These are upper bound values, as they do not include point-source subtraction. We also . give an explanation of the technical difficulties encountered in reducing the data, along with steps taken to resolve them. This will be 10 helpful for other scientific projects which will aim to make use of standalone NenuFAR imaging observations in the future.

11 Keywords: galaxy clusters ;observational cosmology ;radio interferometry ;nenufar

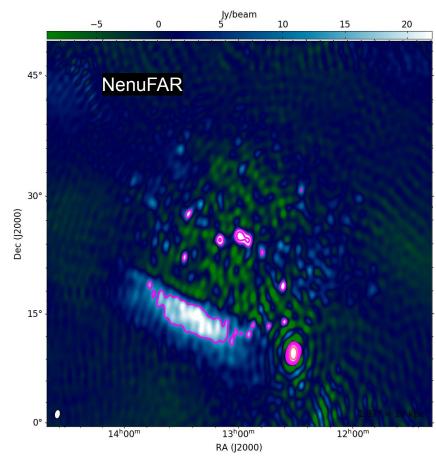
1. Introduction

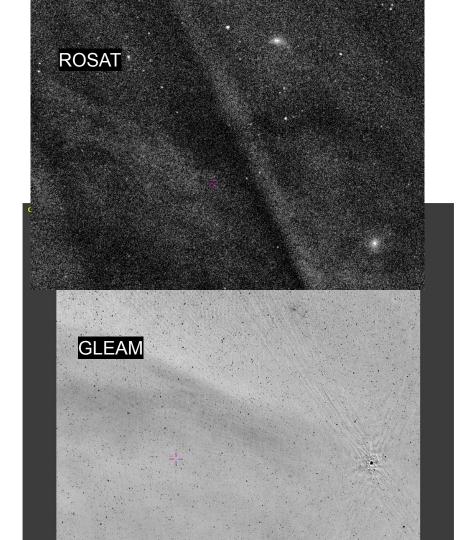
The Coma galaxy cluster is not only one of the first historically observed galaxy clusters [1], but also the first galaxy cluster with the detection of either a radio halo [2] or radio relic [3,4] in the literature. As an object known to host several 26 physical components of considerable scientific interest, it was chosen as the first galaxy cluster to be observed as part of 17 the NenuFAR Cosmic Filaments & Magnetism Pilot Survey project.

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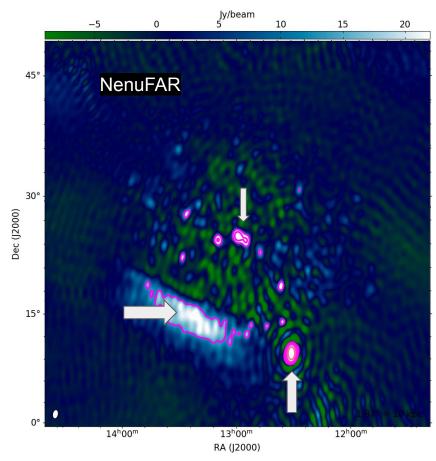
https://www.mdpi.com/journal/galaxies

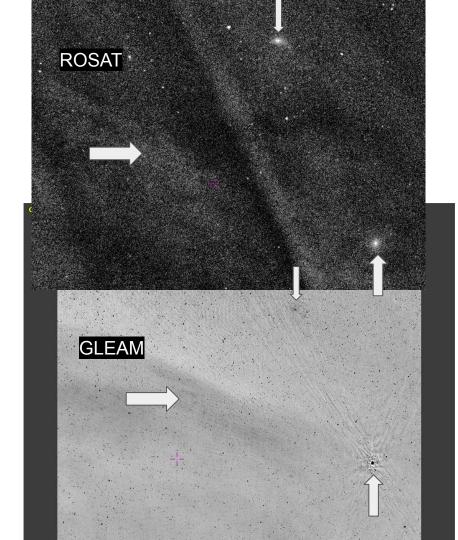
Current Results





Current Results





The Team

KP currently has the following participants:





Etienne Bonnassieux, Annalisa Bonafede, Franco Vazza, Chiara Ferrari,

Francesco de Gasparin, Cyril Tasse, ...









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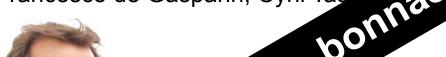
And maybe...you? ;)

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And maybe...you?;)